## IIIITiter

# PROBING NEW LIGHT FORCEMEDIATORS BY ISOTOPE SHIFT 

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Table-Top Experiments with Skyscraper Reach
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[^0]
the Standard Model (SM) works great but it is not a complete picture

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New Physics (NP) is required but its scale is unknown

## THE QUEST FOR NEW PHYSICS

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( TeV scale)

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(up to MeV )

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$$
(\text { up to } \mathrm{MeV})
$$

compare theory to
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## THE QUEST FOR NEW PHYSICS

energy frontier
(TeV scale)
intensity frontier
( MeV - GeV scale)
precision measurements
compare theory to experiment
observables which are insensitive to theory error

## THE QUEST FOR NEW PHYSICS

energy frontier
( TeV scale)
intensity frontier
( MeV - GeV scale)
precision measurements
compare theory to experiment
(up to MeV)
hydrogen and helium


## PRECISION SPECTROSCOPY

## Ytterbium ( $\mathrm{Yb}^{+}$)



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## Ytterbium ( $\mathrm{Yb}^{+}$)



# experimental error of E3 0.25 Hz <br> relative error: $4 \times 10^{-16}$ 

Huntenmann et al. 2014 Gouda et al. 2014
in principle: $y_{e} y_{n}\left(\frac{125 \mathrm{GeV}}{m_{\phi}}\right)^{2}<4 \times 10^{-6}$
stronger than LHC current bounds

## PRECISION SPECTROSCOPY

## Ytterbium ( $\mathrm{Yb}^{+}$)


experimental error
of E3 0.25 Hz
relative error: $4 \times 10^{-16}$
Huntenmann et al. 2014 Gouda et al. 2014
theory is not good enough

## Isotope Shift

## ISOTOPE SHIFT - KING PLOT

the same electronic transition, $i$, in two isotopes, $A$ and $A^{\prime}$

$$
\nu_{i}^{A A^{\prime}} \equiv \nu_{i}^{A}-\nu_{i}^{A^{\prime}}
$$

## ISOTOPE SHIFT - KING PLOT

the same electronic transition, $i$, in two isotopes, $A$ and $A^{\prime}$

$$
\begin{array}{ccc}
\nu_{i}^{A A^{\prime}} \equiv \nu_{i}^{A}-\nu_{i}^{A^{\prime}}=K_{i} \mu_{A A^{\prime}}+F_{i} \delta\left\langle r^{2}\right\rangle_{A A^{\prime}}+\ldots \\
\mu_{A A^{\prime}} \equiv \frac{1}{m_{A}}-\frac{1}{m_{A^{\prime}}} & \text { electronic } & \text { nucleus } \\
\text { parameters } & \text { parameters }
\end{array}
$$

## ISOTOPE SHIFT - KING PLOT

the same electronic transition, $i$, in two isotopes, $A$ and $A^{\prime}$

$$
\left.\begin{array}{ccc} 
& \begin{array}{c}
\text { Mass Shift }
\end{array} & \begin{array}{c}
\text { Field Shift } \\
\text { (short distance) }
\end{array} \\
\nu_{i}^{A A^{\prime}} \equiv \nu_{i}^{A}-\nu_{i}^{A^{\prime}}=K_{i} \mu_{A A^{\prime}}+ & F_{i} \delta\left\langle r^{2}\right\rangle_{A A^{\prime}}+\ldots \\
\mu_{A A^{\prime}} \equiv \frac{1}{m_{A}}-\frac{1}{m_{A^{\prime}}} & \begin{array}{c}
\text { electronic } \\
\text { parameters }
\end{array} & \text { nucleus } \\
m \nu_{i}^{A A^{\prime}} \equiv \nu_{i}^{A A^{\prime}} / \mu_{A A^{\prime}} & \text { parameters } \\
F_{21} \equiv F_{2} / F_{1} \\
K_{21} \equiv K_{2}-F_{21} K_{1} \\
i=1,2
\end{array}\right) m \nu_{2}^{A A^{\prime}}=K_{21}+F_{21} m \nu_{1}^{A A^{\prime}} .
$$

## ISOTOPE SHIFT - KING PLOT

$$
\begin{aligned}
& \overrightarrow{m \nu}_{i} \equiv\left(m \nu_{i}^{A A_{1}^{\prime}}, m \nu_{i}^{A A_{2}^{\prime}}, m \nu_{i}^{A A_{3}^{\prime}}\right) \\
& \overrightarrow{m \mu} \equiv(1,1,1)
\end{aligned}
$$

$$
\overrightarrow{m \nu}_{i}=K_{i} \overrightarrow{m \mu}+F_{i} \overrightarrow{m \delta\left\langle r^{2}\right\rangle}
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\overrightarrow{m \mu} \equiv(1,1,1)
\end{array}
$$



## testing factorization only by data

## ISOTOPE SHIFT - KING PLOT

## existing isotope shift measurement of $\mathrm{Ca}^{+}$



## 100 kHz error

$$
4 \mathrm{~S} \rightarrow 4 \mathrm{P}_{1} / 2
$$

Gebert et al. 2015

## the



## ISOTOPE SHIFT AND NEW PHYSICS

$$
\nu_{i}^{A A^{\prime}}=K_{i} \mu_{A A^{\prime}}+F_{i} \delta\left\langle r^{2}\right\rangle_{A A^{\prime}}+\alpha_{\mathrm{NP}} X_{i} \gamma_{A A^{\prime}}
$$

new physics

## ISOTOPE SHIFT AND NEW PHYSICS

$$
\begin{array}{ll}
\begin{array}{l}
F_{21} \\
\hline \overrightarrow{m \mu} \\
\equiv F_{2} / F_{1} \\
h_{A A^{\prime}} \\
\equiv \gamma_{A A^{\prime}} / \mu_{A A^{\prime}} \\
\overrightarrow{m \nu}_{i}
\end{array}>\left(m \nu_{i}^{A A_{1}^{\prime}}, m \nu_{i}^{A A_{2}^{\prime}}, m \nu_{i}^{A A_{3}^{\prime}}\right)
\end{array} \quad \overrightarrow{m \nu_{2}}=K_{21} \overrightarrow{m \mu}+F_{21} \overrightarrow{m \nu}_{1}+\alpha_{\mathrm{NP}} \vec{h}\left(X_{2}-X_{1} F_{21}\right)
$$



## ISOTOPE SHIFT AND NEW PHYSICS

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\begin{array}{ll}
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\overrightarrow{m \mu} & \equiv(1,1,1) \\
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\overrightarrow{m \nu}_{i} & \equiv\left(m \nu_{i}^{A A_{1}^{\prime}}, m \nu_{i}^{A A_{2}^{\prime}}, m \nu_{i}^{A A_{3}^{\prime}}\right)
\end{array}
$$



- $X_{2} \neq X_{1} F_{21}$ - long distance NP
- $h$ - is not aligned with $m \nu_{1}, m \nu_{2}, m \mu$
nonlinear King plot from NP


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nonlinear King plot from NP

$$
\alpha_{\mathrm{NP}}=\frac{\left(\overrightarrow{m \nu}_{1} \times \overrightarrow{m \nu}_{2}\right) \cdot \overrightarrow{m \mu}}{(\overrightarrow{m \mu} \times \vec{h}) \cdot\left(X_{1} \overrightarrow{m \nu}_{2}-X_{2} \overrightarrow{m \nu}_{1}\right)}
$$

the only theory inputs
similar to data driven background estimation at the LHC

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the only theory inputs
similar to data driven background estimation at the LHC
data consistent with linearity

## CONSTRAINING LIGHT NEW BOSONS

new bosons with couplings to $e$ and $n$
(spin independent)

$$
V_{\phi}(r)=\alpha_{\mathrm{NP}}(A-Z) \frac{e^{-m_{\phi} r}}{r}
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## CONSTRAINING LIGHT NEW BOSONS

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$V_{\phi}(r)=\alpha_{\mathrm{NP}}(A-Z) \frac{e^{-m_{\phi} r}}{r}$
$X_{i}=\int d^{3} r \frac{e^{-m_{\phi} r}}{r}\left[\left|\Psi_{b}(r)\right|^{2}-\left|\Psi_{a}(r)\right|^{2}\right]$
$1^{\text {st }}$ order perturbation theory and multi-body perturbation theory

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$$
V_{\phi}(r)=\alpha_{\mathrm{NP}}(A-Z) \frac{e^{-m_{\phi} r}}{r}
$$

$$
\alpha_{\mathrm{NP}}=\frac{y_{e} y_{n}}{4 \pi} \quad h_{A A^{\prime}} \propto A A^{\prime}
$$

$X_{i}=\int d^{3} r \frac{e^{-m_{\phi} r}}{r}\left[\left|\Psi_{b}(r)\right|^{2}-\left|\Psi_{a}(r)\right|^{2}\right]$
$1^{\text {st }}$ order perturbation theory and multi-body perturbation theory

|  | $m_{\phi}<4 \mathrm{keV}$ | $4 \mathrm{keV}<m_{\phi}<50 \mathrm{MeV}$ | $50 \mathrm{MeV}<m_{\phi}$ |
| :---: | :---: | :---: | :---: |
| $\mathrm{V}_{\phi}(r) \sim$ | $1 / r$ | $\exp \left(-m_{\phi} r\right) / r$ | $\delta(r) /\left(m_{\phi} r\right)^{2}$ |
| $X_{i}$ | constant | $m_{\phi}$ dependent | $X_{2}-X_{1} F_{21} \rightarrow 0$ |

## CONSTRAINING LIGHT NEW BOSONS



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## CONSTRAINING LIGHT NEW BOSONS



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- current data:
- Ca $^{+}: 866 / 397 \mathrm{~nm}, \sigma \sim 0.1 \mathrm{MHz}$
- $\mathbf{Y b} \mathbf{b}^{0}: 556 / 399 \mathrm{~nm}, \sigma \sim 0.1-0.5 \mathrm{MHz}$


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- current data:
- Ca ${ }^{+}: 866 / 397 \mathrm{~nm}, \sigma \sim 0.1 \mathrm{MHz}$
- $\mathbf{Y b}^{0}: 556 / 399 \mathrm{~nm}, \sigma \sim 0.1-0.5 \mathrm{MHz}$
- candidates for future measurements:
- $\mathrm{Ca}^{+}: S \rightarrow D_{5 / 2} / S \rightarrow D_{3 / 2}$
- $\mathrm{Sr}^{+}: S \rightarrow \mathrm{D}_{5 / 2} / S \rightarrow \mathrm{D}_{3 / 2}$
- $\mathbf{S r}^{+} / \mathbf{S r}: S \rightarrow P / S \rightarrow D_{5 / 2}$
- $\mathbf{Y} \mathbf{b}^{+}: S \rightarrow D_{3 / 2} / S \rightarrow F_{7 / 2}$


## BOUNDSAND PROJECTIONS



## BOUNDS AND PROJECTIONS


few electrons atoms

# HYDROGEN AND HELIUM SPECTROSCOPY 

## direct comparison of theory to experiment (not limited by theory error)

bound Yukawa like force with spin independent interactions:

$$
\begin{gathered}
\frac{y_{e}\left(y_{p} Z+(A-Z) y_{n}\right)}{4 \pi} \frac{e^{-m_{\phi} r}}{r} \\
\frac{y_{e}^{2}}{4 \pi} \frac{e^{-m_{\phi} r_{12}}}{r_{12}}
\end{gathered}
$$

## HYDROGEN AND HELIUM SPECTROSCOPY

## direct comparison of theory to experiment (not limited by theory error)

bound Yukawa like force with spin independent interactions:

## hydrogen

helium
isotope shift (He3-He4, H-D)
positronium
$\frac{y_{e}\left(y_{p} Z+(A-Z) y_{n}\right)}{4 \pi} \frac{e^{-m_{\phi} r}}{r}$

$$
\frac{y_{e}^{2}}{4 \pi} \frac{e^{-m_{\phi} r_{12}}}{r_{12}}
$$


$y_{e} y_{n}$
ye

## HYDROGEN AND HELIUM SPECTROSCOPY

## isotope shift



## HYDROGEN AND HELIUM SPECTROSCOPY

electron interaction


## SUMMARY



## SUMMARY



- precision isotope spectroscopy can probe new light force-carriers with spin independent couplings to the electron and neutron
- King analysis has minimal theory inputs ("data-driven background")
- current constraints from King analysis are weak - but future measurements may improve the state-of-the-art bounds

BACKUP SLIDES

## be ANomALY



Frugiuele, Fuchs, Perez, Schlaffer - 1602.04822


[^0]:    C. Delaunay, C. Frugiuele, E. Fuchs, YS - work in progress
    J.C. Berengut, D. Budker, C. Delaunay, V.V. Flambaum, C. Frugiuele, E. Fuchs, C. Grojean, R. Harnik, R. Ozeri, G. Perez, YS - 1704.06005
    C. Delaunay, R. Ozeri, G. Perez, YS 1601.05087

